

HOW TO CREATE A PERSPECTIVE PROJECTION IN IDL

Do the following:

1. Decide where you want the pole of your projection to be. You can define the center in terms of Galactic or equatorial coordinates; we'll talk in terms of Galactic coordinates here as an example. Let the Galactic coordinates of your perspective pole be (gl_{pole}, bg_{pole}) .

Your projection pole defines a new spherical coordinate system. A particular position in Galactic coordinates is specified by (gl, gb) and in your new system it is specified by the angles $(long, lat)$. Given (gl, gb) , you must find $(long, lat)$ using a proper spherical coordinate transformation using the rotation matrix **rtot** that relates your projection to the Galactic one.

To determine the rotation matrix **rtot**, use our procedure

`rmatrix, gl_pole, gb_pole, rtot`

2. For your particular point, we have our coordinate conversion procedure

`sph_coord_conv, gl, gb, rtot, long, lat`

3. The perspective projection represents the your particular position's angular coordinates $(long, lat)$ in terms of linear coordinates on the projection plane x and y , whose origin is the pole of your projection. The origin has sky coordinates (gl_{pole}, bg_{pole}) . It's conceptually simpler to think of the projection plane in terms of polar coordinates R and θ , where as usual $R = (x^2 + y^2)^{1/2}$ and $\theta = \text{atan}(y/x)$. Then R is uniquely related to lat and θ to $long$.

For a stereographic projection,

$$R = \tan\left(\frac{90. - lat}{2}\right) \quad (1)$$

and

$$\theta = long + longoffset \quad (2)$$

where *longoffset* is a zero point of your own choosing and is a keyword in **rmatrix**; you can leave it equal to zero. For a gnomonic projection, replace equation 1 by

$$R = \tan(90. - lat) \quad (3)$$

For these, you can use our procedure **stereographic.pro** .